

X-641-71-481

PREPRINT

NASA TM X

65781

G3/19

(NASA-TM-X-65781) ON LASHINSKY'S EQUATION
y'' PLUS alpha y' MINUS ay PLUS by CUBED
EQUALS 0 F.F. Cap (NASA) Dec. 1971 16 p
CSCL 12A

N72-14620

Unclass
11358

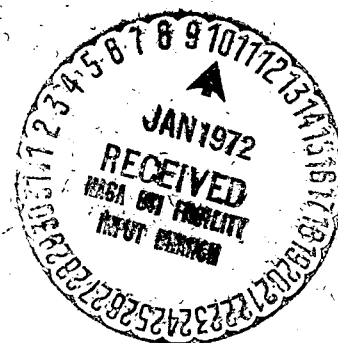
U.S. (NASA) CR 12A

ON LASHINSKY'S EQUATION

$$y'' + \alpha y' - ay + by^3 = 0$$

FERDINAND F. CAP

DECEMBER 1971



GODDARD SPACE FLIGHT CENTER

GREENBELT, MARYLAND

Reproduced by
NATIONAL TECHNICAL
INFORMATION SERVICE
U.S. Department of Commerce
Springfield VA 22151

INTRODUCTION

The equation

$$y'' + \alpha y' - ay + by^3 = 0 \quad (1)$$

has been proposed by H. Lashinsky¹ as an analog to the Van der Pol equation² to describe nonlinear saturation of aperiodic instabilities such as the Benard instability in fluids³ and related plasma instabilities⁴.

In the Van der Pol equation, the saturation is described by a nonlinear dissipation term of the form $(-\alpha + \beta y + \gamma y^2)y'$. The instability corresponds to the negative sign of the coefficient α : This negative dissipation feeds energy into the system. The nonlinear terms saturate the instability.

In the Lashinsky equation, on the other hand, the saturation is described by a nonlinear restoring force of the form $(-a + by^2)y$. The instability itself corresponds to the negative sign of the coefficient a in the restoring force. The dissipation term consumes energy in this case.

OSCILLATORY SOLUTIONS

First we wish to determine whether equation (1) has periodic or oscillatory solutions for $a > 0$, $\alpha > 0$, $b > 0$. The equation (1) is a special form of the more general Liénard equation

$$y'' + f(y) \cdot y' + g(y) = 0 \quad (2)$$

which has been investigated extensively^{5,6}.

The Liénard Theorem^{6,7,8} gives sufficient but not necessary conditions for the existence of a strictly periodic solution $y(t)$. The conditions to be satisfied are:

1. $f(y) = \alpha$, $g(y) = -ay + by^3$ analytic, which is satisfied for (1).
2. $g(y)$ an odd function, $g(0) = 0$ and $yg(y) \geq 0$, so that a restoring force exists. This condition is also satisfied, depending on the values of the coefficients a and b .
3. $f(y)$ an even function and $f(0) < 0$ (negative damping at the origin, so that there is an unstable singular point at the origin). For (1) $f(0)$ is positive, however, and the damping term is not a source of energy.
4. $F(y) = \int_0^y f(y)dy = \alpha y$, $G(y) = \int_0^y g(y)dy = -\frac{a}{2}y^2 + b\frac{y^4}{4}$ go to ∞ as $y \rightarrow \infty$, which insures periodicity. This condition is satisfied.
5. $F(0) = 0$ has a unique root at y_0 , $y_0 > 0$ and is a monotonically increasing function for $y = y_0$. Since we have $y_0 = 0$, the condition $y_0 > 0$ is not satisfied.

Since all the conditions for periodicity are not satisfied, the Lashinsky equation cannot possess a periodic solution. Also the Bendixson Theorem⁵, the Frommer Theorem⁸, the Levinson-Smith Theorem^{5,6} show that (1) has no periodic solutions in the sense

$$y(t + t_0) = y(t) \quad (3)$$

This result, however, does not preclude the existence of oscillatory solutions, i.e. solutions with several maxima and minima. Apparently these oscillatory solutions are damped.

In order to investigate these solutions, we use the phase plane method⁸. We write (1) in the form of two first order equations

$$\begin{aligned} y' &= v \equiv P(v, y) \\ v' &= -\alpha v + ay - by^3 \equiv Q(v, y) \end{aligned} \quad (4)$$

Critical (equilibrium) points are then given by $y' = 0$, $v' = 0$; these points are i.e.

$$y' = v = 0 \quad y = \pm \sqrt{\frac{a}{b}}, 0 \quad (5)$$

so that $y'' = 0$ at these points. The characteristic equation that determines the nature of these critical points is then

$$\lambda^2 - (P_y + Q_v)\lambda + P_y Q_v - P_v Q_y = 0 \quad (6)$$

From (4) we have $P_y = 0$, $P_v = 1$, $Q_y = a - 3by^2$, $Q_v = -\alpha$, so that (6) becomes

$$\lambda^2 + \alpha\lambda - a + 3by^2 = 0 \quad (7)$$

which gives for the point $y = 0$, $v = 0$

$$\lambda_{1,2} = -\frac{\alpha}{2} \pm \sqrt{\frac{\alpha^2}{4} + a} \quad (8)$$

so that the origin of the phase plane is a saddle point (two real roots λ with opposite signs), so that this solution is unstable and $y \rightarrow \infty$,

$u \rightarrow \infty$ for $t \rightarrow \infty$. The critical points $y = \pm \sqrt{\frac{a}{b}}$, $u = 0$ give

$$\lambda_{1,2} = -\frac{\alpha}{2} \pm \sqrt{\frac{\alpha^2}{4} - 2a} \quad (9)$$

so that these points are either a stable focus (complex $\lambda_{1,2}$ with negative real parts) for $\alpha^2 < 8a$, indicating a damped oscillation, or a stable mode (both λ are real and negative) for $\alpha^2 > 8a$, indicating that no oscillation occurs and $y \rightarrow \text{const}$ or zero, $u \rightarrow 0$ for $t \rightarrow \infty$.

In the case given by Lashinsky¹, $\alpha \approx 10$, $a \approx 1$ and we have no oscillating solution, see section 3. In this section we investigate oscillatory solutions, assuming $\alpha \approx 0$, $1, a \approx 1$. It is of interest that the value of b does not enter directly into this discussion which is based on a linearized stability analysis. This feature is noteworthy, since both terms $\alpha y'$ and $-ay$ are nonoscillatory and the only term that produces oscillations is by^3 . The coefficient b enters the discussion only through (5).

In order to eliminate b completely, we make the transformation

$$y(x) = a^{1/2} b^{-1/2} u(t) \quad (10)$$

which gives

$$u'' + \alpha u' - au + au^3 = 0 \quad (11)$$

In the oscillatory case $\alpha^2 < 8a$ and the smaller α the weaker the dissipation. Neglecting $\alpha u'$ as a first approximation, by multiplying (11)

by u' and integrating we find

$$\frac{u'^2}{2} - \frac{au^2}{2} + \frac{au^4}{4} = E \quad (12)$$

where E is an integration constant (the energy). In this conservative system no energy is lost and energy oscillates between kinetic and potential energy. If a small dissipative term $\alpha u'$ is now introduced; then E decreases slowly and $E = \text{const}$ must be replaced by $E(t)$. This procedure can be used to establish a perturbation method based on the averaging process⁹.

From (12) we have

$$t(u) + \varphi = \int \frac{du}{\sqrt{2E + au^2 - \frac{au^4}{2}}} \quad (13)$$

where φ is another integration constant. The integral is an elliptic integral and may be solved by standard methods¹⁰.

The inverse function $u(t)$ - and also $y(t)$ - may be expressed by Jacobi elliptic functions whose modulus k depends on the energy constant $E^{11,12}$. If E varies slowly because of the dissipative term $\alpha u'$, then $k \rightarrow k(t)$ and we can use the averaging method we have described elsewhere⁹.

NON-OSCILLATORY SOLUTIONS

This is the case discussed by H. Lashinsky¹. For $\alpha \approx 10$, $a \approx 1$ the non-oscillatory condition $\alpha^2 > 8a$ is satisfied. We start from (11) using the transformation $\alpha t = \xi$, so that

$$\ddot{u} + \dot{u} - \varepsilon u + \varepsilon u^3 = 0 \quad (14)$$

where $\dot{u} = du/d\xi$ and $\varepsilon = a/\alpha^2 = 10^{-2}$. Now the perturbation problem is no longer singular (as it was for $\alpha^2 < 8a$) and we may set up a Poincare expansion of successive approximations

$$u(\xi) = u_0(\xi) + \varepsilon u_1(\xi) + \varepsilon^2 u_2(\xi) + \dots \quad (15)$$

Substituting into (14) we obtain

$$\ddot{u}_0 + \dot{u}_0 = 0 \quad (16)$$

$$\ddot{u}_1 + \dot{u}_1 = u_0 - u_0^3 \quad (17)$$

$$\ddot{u}_2 + \dot{u}_2 = u_1 - 3u_0^2 u_1 \quad (18)$$

neglecting higher orders than ε^2 . Integration yields

$$u_0(\xi) = Ae^{-\xi} + B \quad (19)$$

$$u_1(\xi) = -A\xi e^{-\xi} - A^3 \frac{1}{6} e^{-3\xi} - \frac{3A^2B}{2} e^{-2\xi} + 3AB^2\xi e^{-\xi} - (-B+B^3)\xi + Ce^{-\xi} + D \quad (20)$$

For $\xi \rightarrow \infty$, the terms containing $B\xi$, $B^3\xi$ tend to infinity, so that the expansion (15) does not converge for a finite number of terms. These terms are typical secular terms. In the oscillatory solution the secular terms are avoided by the averaging process⁹. In the present case the secular terms are eliminated by appropriate adjustment of the arbitrary integration constants (Lindstedt method¹³). To do this we put $B = 0$, $D = 0$. The terms $\xi e^{-\xi}$ etc. are not secular. For $u_2(\xi)$ we have

$$u_2(\xi) = \frac{7}{18} A^3 e^{-3\xi} + \frac{1}{2} A^3 \xi e^{-3\xi} + \frac{1}{40} A^5 e^{-5\xi} + \frac{A}{2} \xi^2 e^{-\xi} + \quad (21)$$

$$A\xi e^{-\xi} - C\xi e^{-\xi} - \frac{A^2C}{2} e^{-3\xi} + Ee^{-\xi} + F$$

For $\xi \rightarrow \infty$ we now have from (15) and (19) - (21)

$$u(\infty) = \mathcal{E}^2 F, \quad u'(\infty) = 0, \quad u''(\infty) = 0 \quad (22)$$

so that from (14) $F = \mathcal{E}^{-2}$ follows.

The following two initial conditions are of physical interest:

$$u(0) = d, \quad \dot{u}(0) = f$$

where

$$\text{a) } f = 0, d \text{ given;} \quad \text{b) } f \text{ given, } d = 0.$$

We then have two equations for the three remaining constants A , C , E so that one constant may be chosen arbitrarily e.g. to give a better fit

to data obtained from a numerical integration. However, in order to adjust a solution of the type (15) to the solution obtained from a numerical integration, higher powers of \mathcal{E} than the second must be taken into account. Thus a straightforward numerical integration is simpler than the expansion (15). The expansion, however, gives more physical insight into the phenomena. The energy fed into the system at $t = 0$ is used to excite several modes - the higher the initial energy, the greater the number of modes that are excited. The higher the order of the mode, the more rapidly it decays, so that after a certain time only the asymptotic value 1 remains.

The result of the numerical integration for $\mathcal{E} = 10^{-2}$, $f = 0$, $d = 0.2$ are shown in Fig. 1; the values are given in Table 1. For $\mathcal{E} = 10^{-2}$, $f = 0.2$, $d = 0$ we find the values in Fig. 2 and Table 2.

ACKNOWLEDGMENT

The author is indebted to the U.S. National Academy of Sciences and the National Research Council for the award of a Senior Postdoctoral Research Associateship; also to NASA-Goddard Space Flight Center for their support and hospitality and for providing a stimulating environment. Especially the author would like to thank Prof. H. Lashinsky, Institute for Fluid Dynamics and Applied Mathematics, University of Maryland, for suggesting the problem and Mr. Peter Smidinger, Goddard Space Flight Center for his help in the numerical integration.

REFERENCES

1. H. Lashinsky, Notes for Seminar in Nonlinear Oscillations, Institute for Fluid Dynamics and Applied Mathematics, University of Maryland, 1970-1971 (to be published).
2. Y. Hale, Oscillations in Nonlinear Systems, McGraw Hill, New York, 1963.
3. F. Cap, Magnetohydrodynamik, Vieweg, Braunschweig, 1972.
4. A. Yedenov, E. Velikhov, R. Sagdeev, Stability of Plasma, Soviet Physics Uspeklu, 4, 2, 332-369 (1961).
5. L. Cesari, Asymptotic Behavior and Stability Problems in Ordinary Differential Equations, Springer, Berlin, 1963.
6. N. Minorsky, Nonlinear Oscillations, v. Nostrand, Princeton, 1962.
7. N. Butenin, Elements of the Theory of Nonlinear Oscillations, Blaisdel , New York, 1965.
8. H. Davis, Introduction to Nonlinear Differential and Integral Equations, Dover, New York, 1962.
9. F. Cap, Averaging Method for the Solution of Nonlinear Differential Equations with Periodic Nonharmonic Solutions, Report X-640-71-396, September 1971, Goddard Space Flight Center, Greenbelt, Maryland 20771 (will be published).
10. W. Groebner, N. Hofreiter, Integraltafel, Springer, Wien, 1949.
11. F. Cap, Bull. Amer. Phys. Soc. (2) 16, 11, 1274 (1971).
12. F. Cap, Langerin Equation of Motion of Electrons in a Nonhomogeneous Plasma, Letter to the Editor, Nuclear Fusion, January 1972.
13. A. Lindstedt, Mém. de l'Ac. Impér. Petersburg, 31, 1883.

Table 1

| \underline{x} | \underline{y} |
|-----------------|-----------------|
| 0 | 0.20000 |
| 0.09999 | 0.20009 |
| 0.49999 | 0.20020 |
| 0.99999 | 0.20070 |
| 4.99986 | 0.20776 |
| 9.99969 | 0.21783 |
| 19.9973 | 0.23928 |
| 39.9863 | 0.28779 |
| 59.9757 | 0.34410 |
| 79.0148 | 0.40493 |
| 100.003 | 0.47911 |
| 120.992 | 0.55836 |
| 126.988 (a) | 0.58143 |
| 139.981 | 0.63125 |
| 160.020 | 0.70539 |
| 180.009 | 0.77259 |
| 200.998 | 0.83272 |
| 219.987 | 0.87670 |
| 250.021 | 0.92699 |
| 280.004 | 0.95840 |
| 309.989 | 0.97684 |
| 343.984 | 0.98959 |
| 375.912 | 0.993722 |
| 400.000 | 0.99621 |

Table 2

| \underline{x} | \underline{y} |
|-----------------|-----------------|
| 0 | 0 |
| 0.04999 | 0.009754 |
| 0.09999 | 0.019033 |
| 0.99999 | 0.126628 |
| 4.99986 | 0.20461 |
| 9.99969 (b) | 0.21584 |
| 19.9973 | 0.23713 |
| 39.863 | 0.285267 |
| 59.9753 | 0.34120 |
| 79.0148 | 0.40169 |
| 100.003 | 0.47556 |
| 120.992 | 0.55466 |
| 127.988 (a) | 0.58156 |
| 139.981 | 0.62758 |
| 160.020 | 0.70149 |
| 180.009 | 0.76956 |
| 200.998 | 0.83022 |
| 219.987 | 0.87471 |
| 250.021 | 0.92570 |
| 280.004 | 0.95763 |
| 309.989 | 0.97639 |
| 349.984 | 0.98939 |
| 375.012 | 0.99360 |
| 400.000 | 0.99614 |

(a) Here we have a maximum of y' .

(b) Near here we have a minimum of y' .

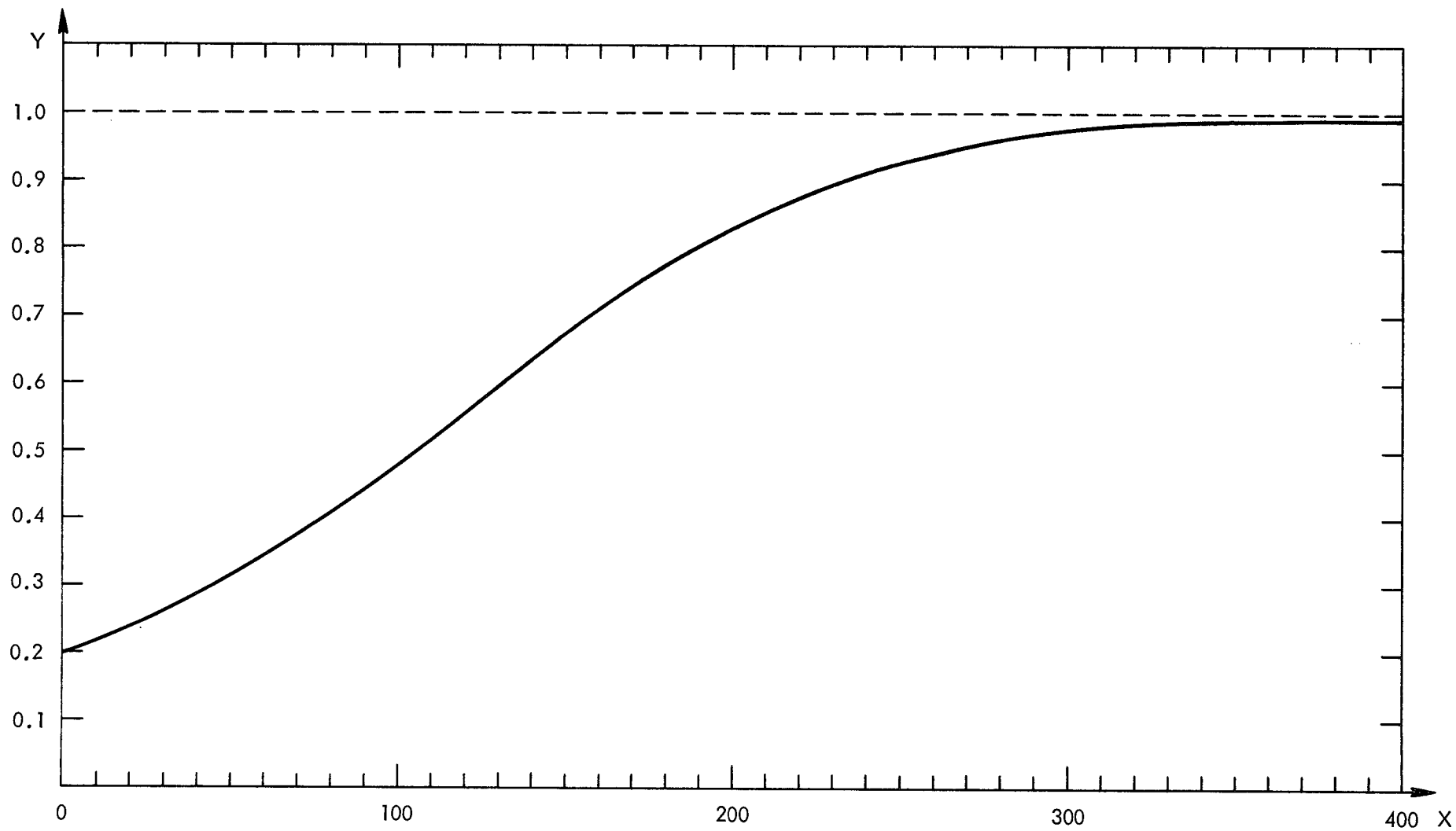


Figure 1.

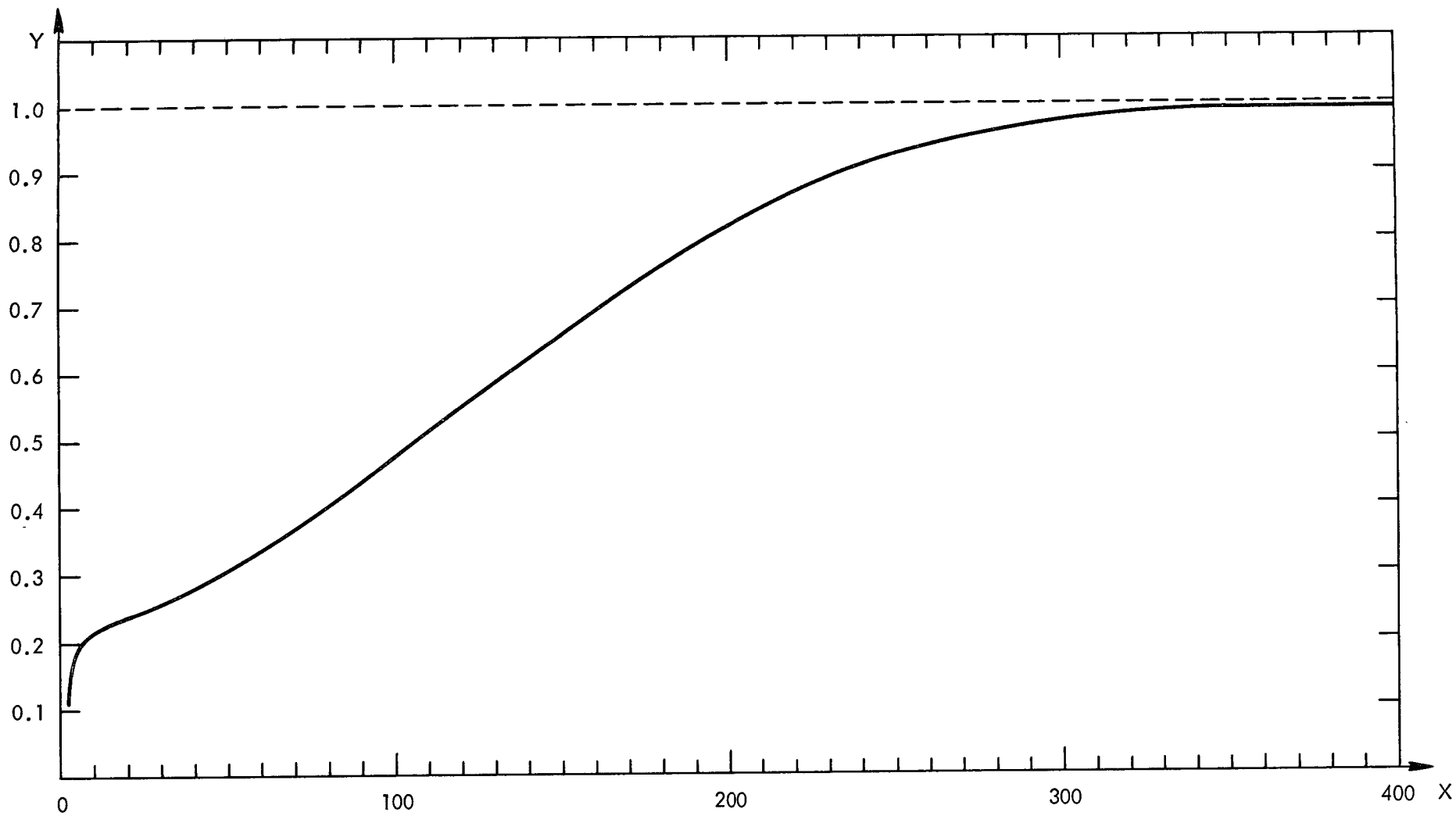


Figure 2.